

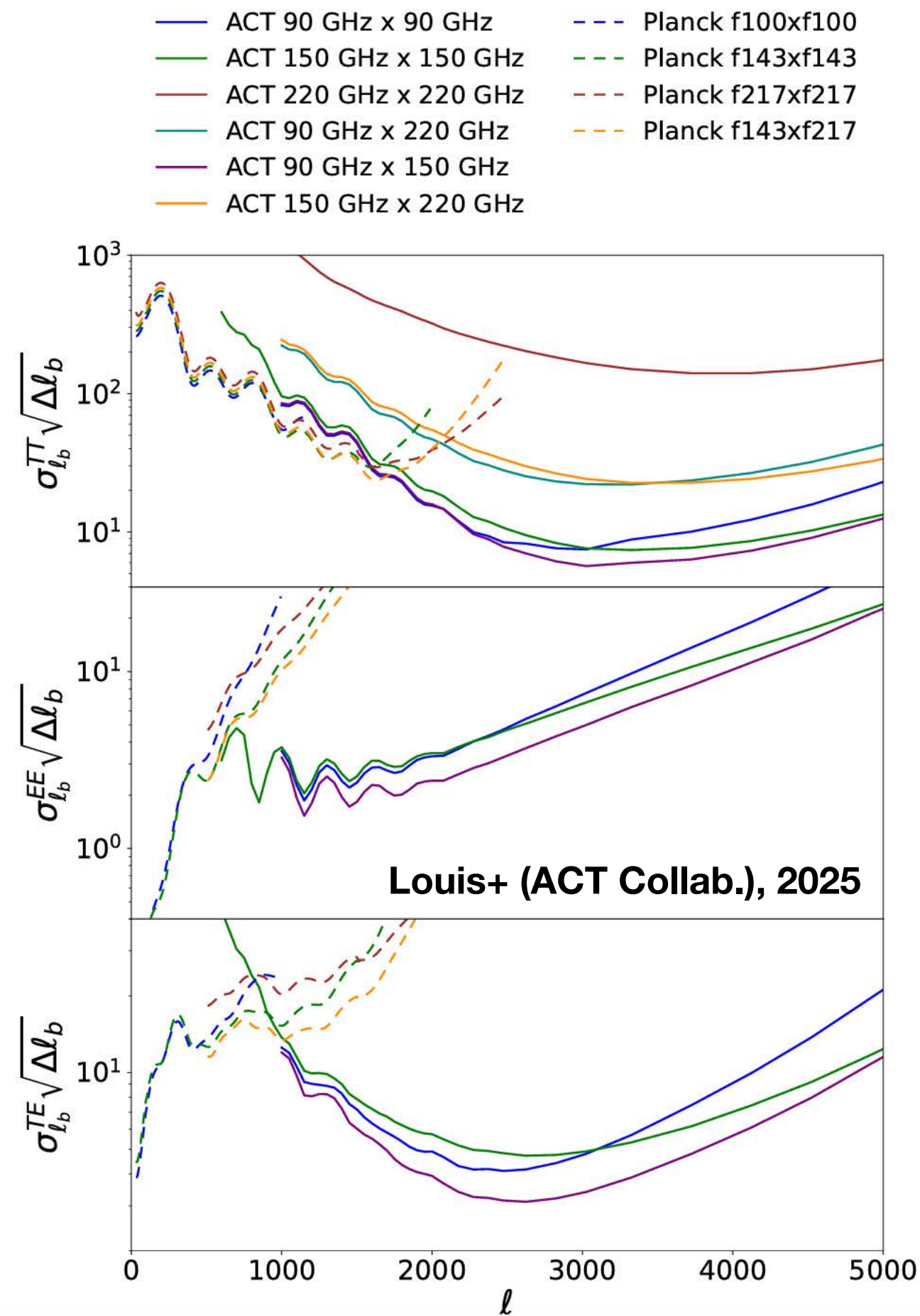
CMB@60

Panel 7 - Space and Ground complementarity

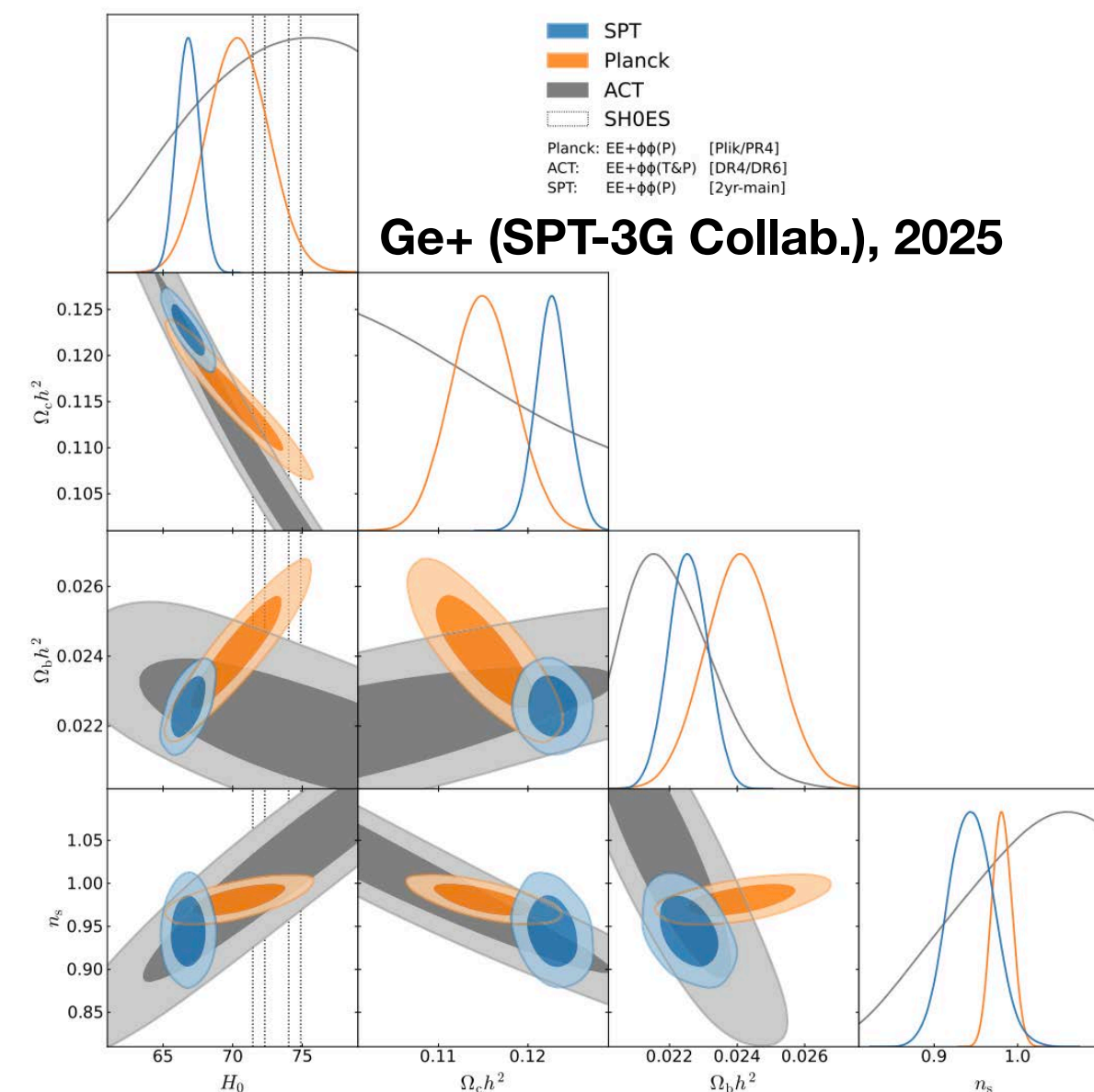
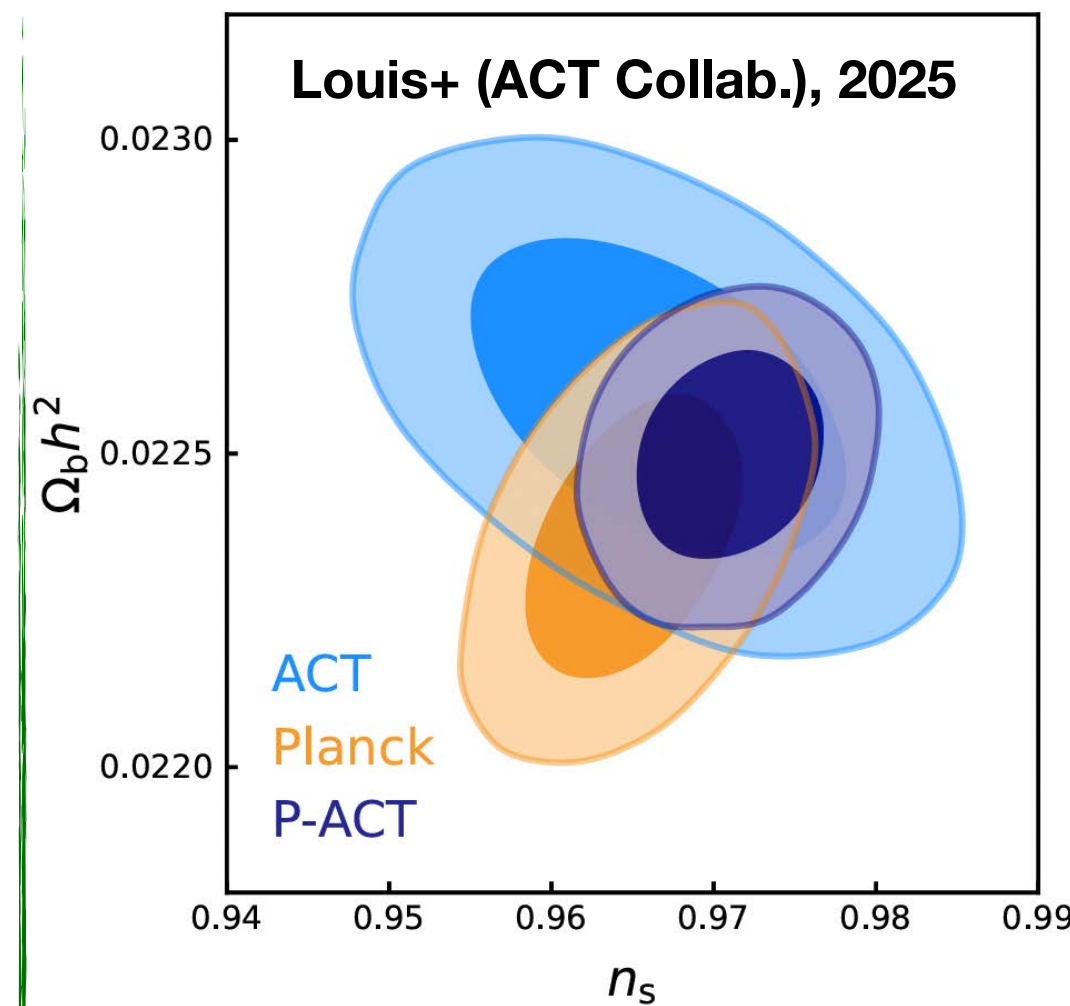
Martina Gerbino - INFN Ferrara



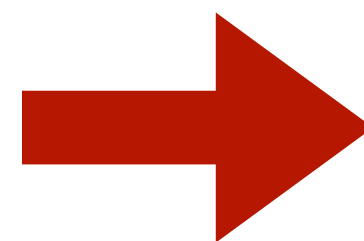
Space vs/+ Ground



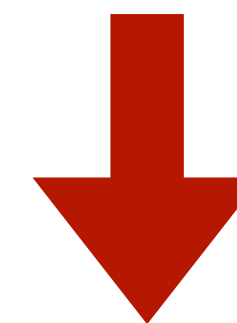
- Access to different scales with different sensitivity



- **Constraining power now comparable**
- **Complementary cosmological information**



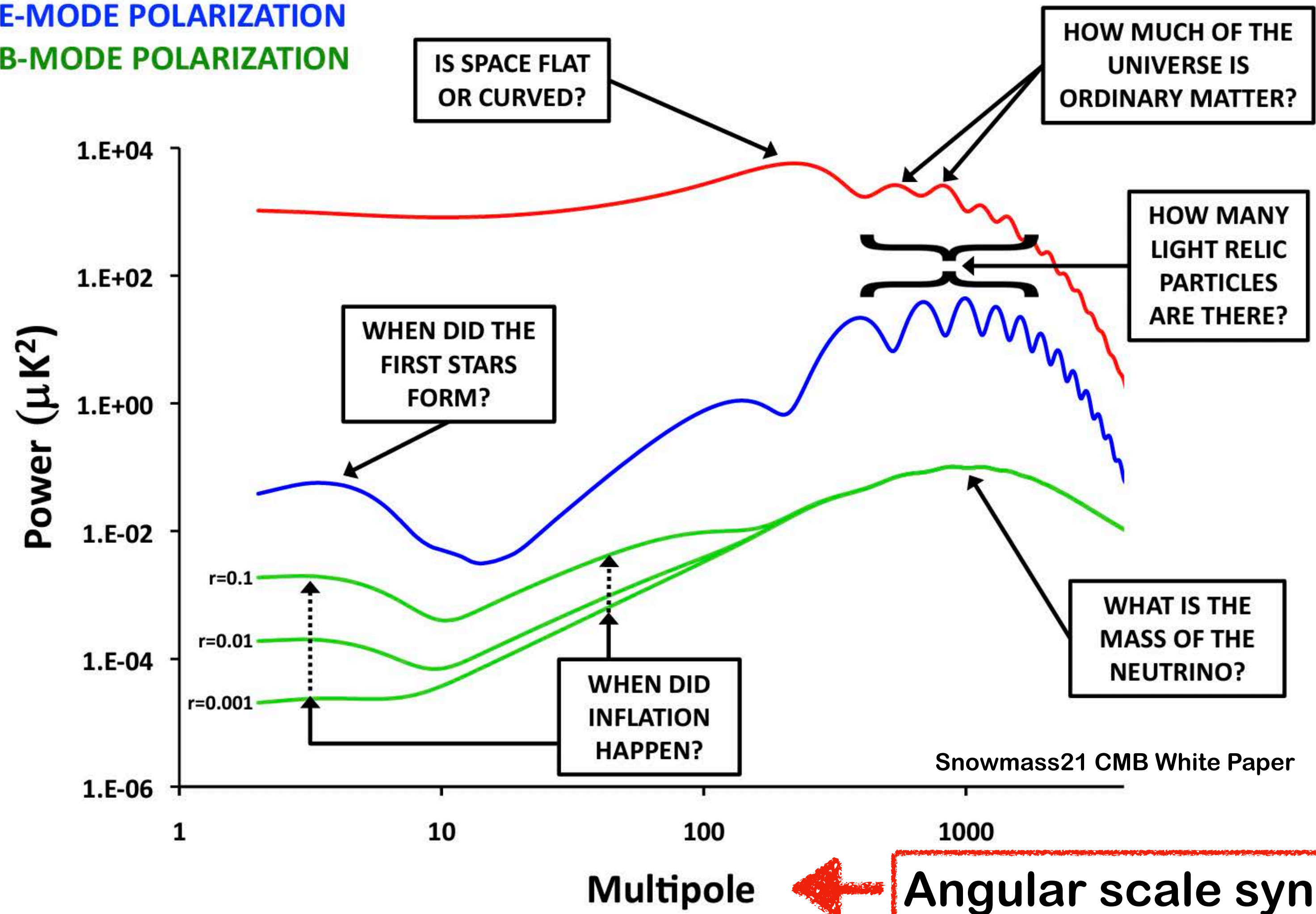
- **Space vs ground = strong cross-check**
- **Space + ground = improved sensitivity**



What from where

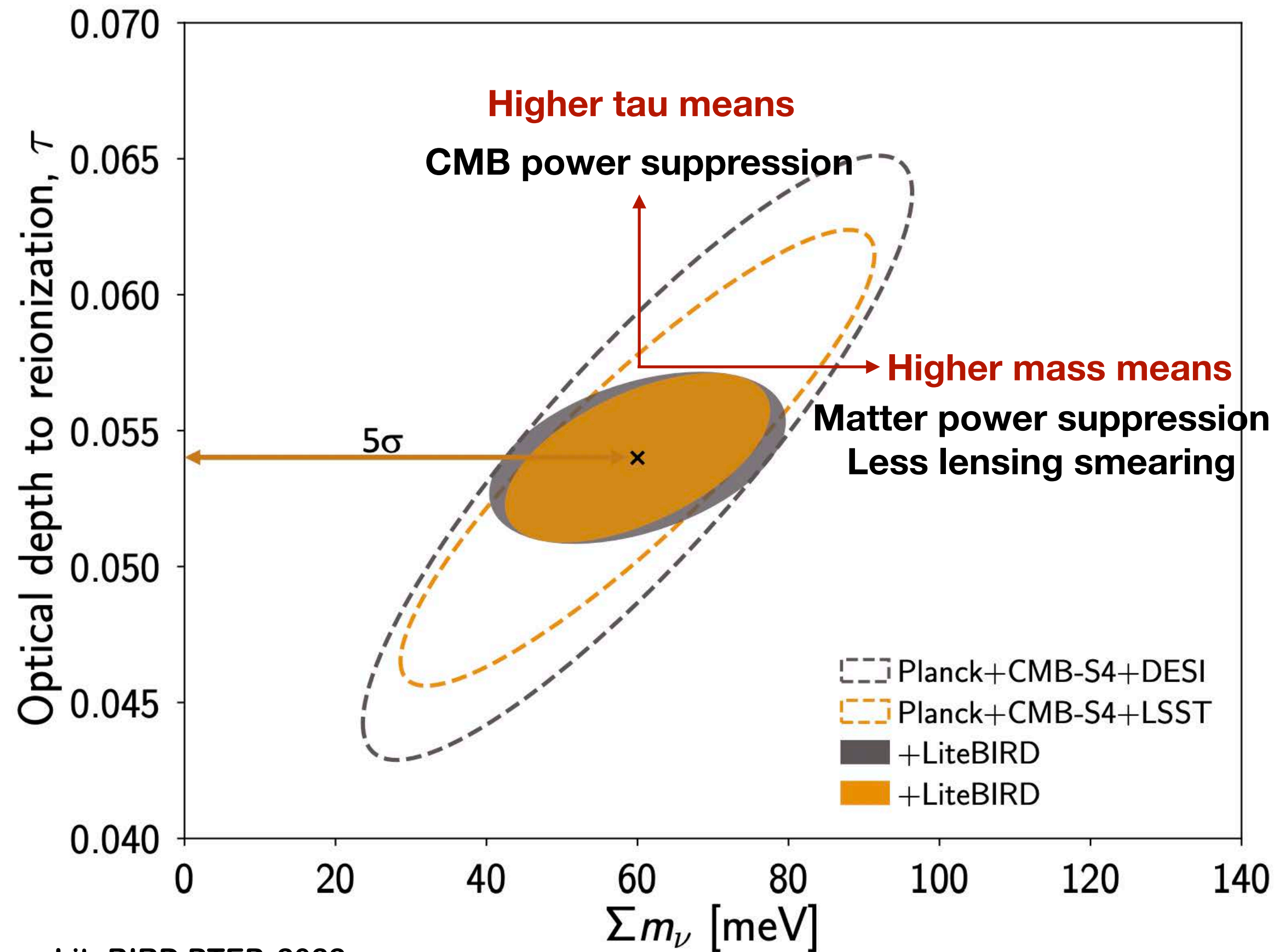
Field synergy

TEMPERATURE
E-MODE POLARIZATION
B-MODE POLARIZATION



Angular scale synergy

Particle cosmology



LiteBIRD PTEP, 2022

See also M. Lattanzi slides
and Panel 4

Measurement of neutrino masses
requires comparison between
high-z and low-z amplitude
from high-res experiments

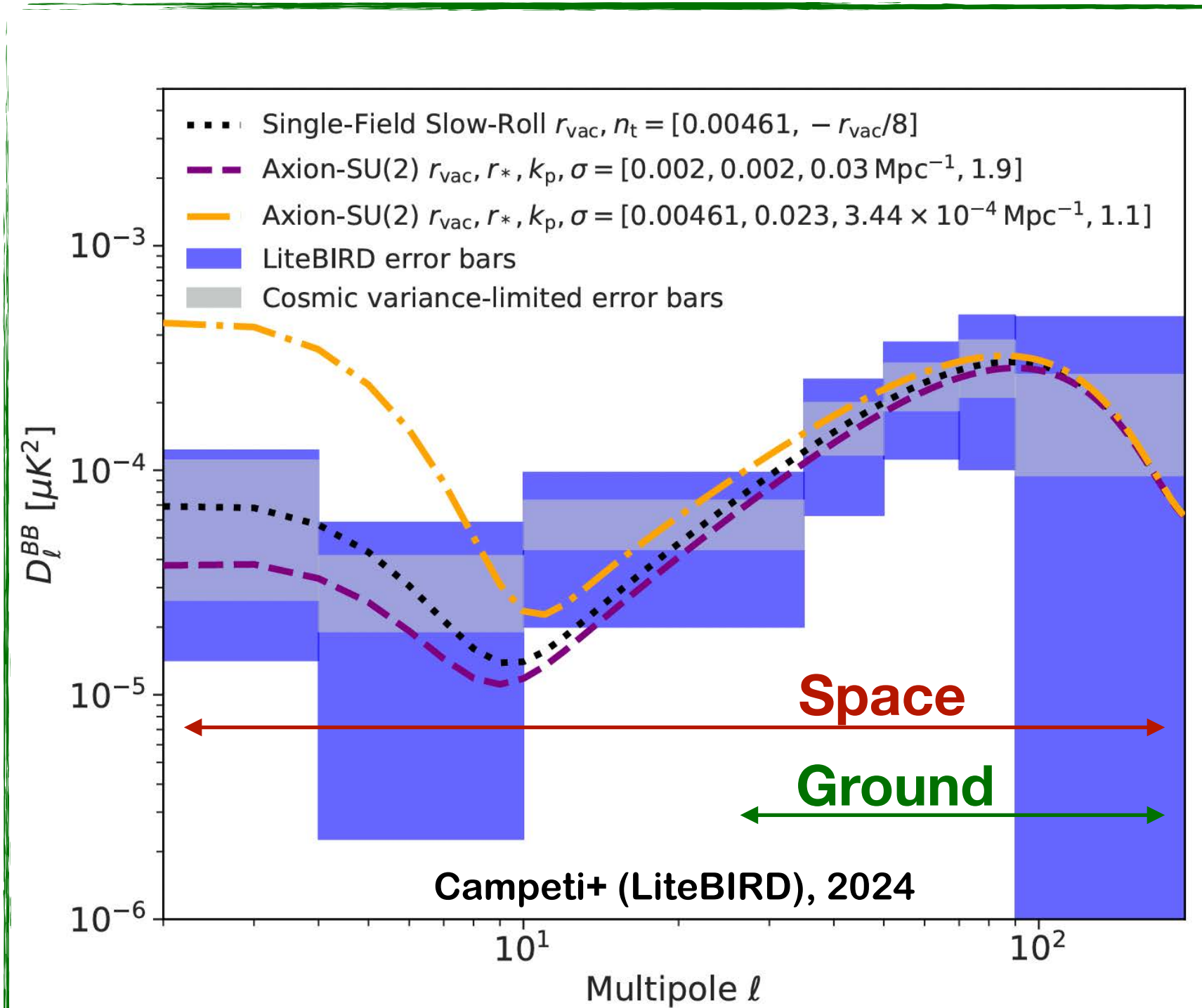
High-z amplitude convolved with tau

Need for CVL measurement of tau
from large-scale polarisation

Combine the power of
large-angle high-sensitivity CMB polarisation
with arcmin maps from ground

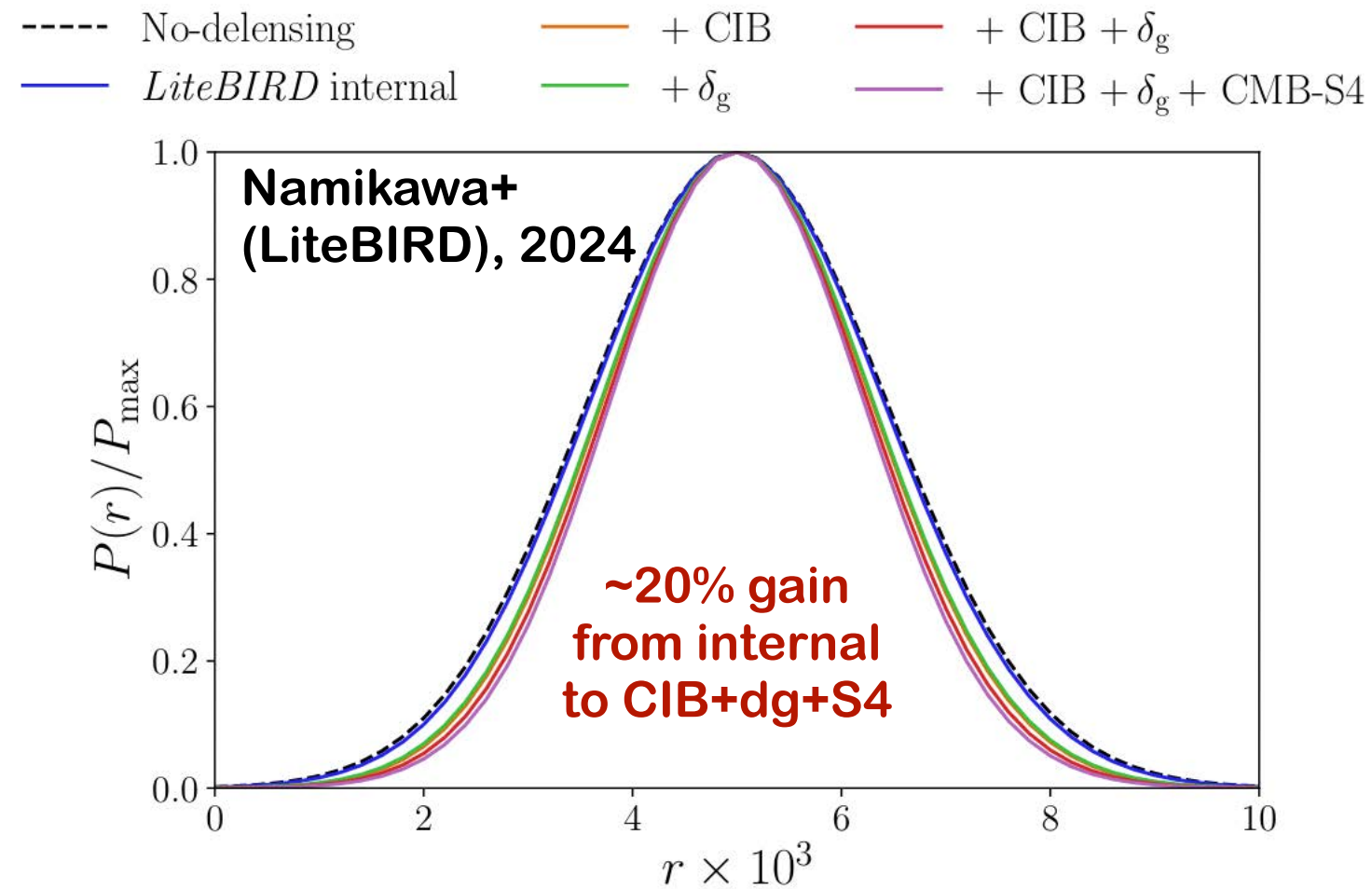
See recent works claiming for a higher tau
to “solve” the negative neutrino mass DESI preference

Early Universe



Tensor modes visible in both reionisation and recombination bumps
 If both detected:

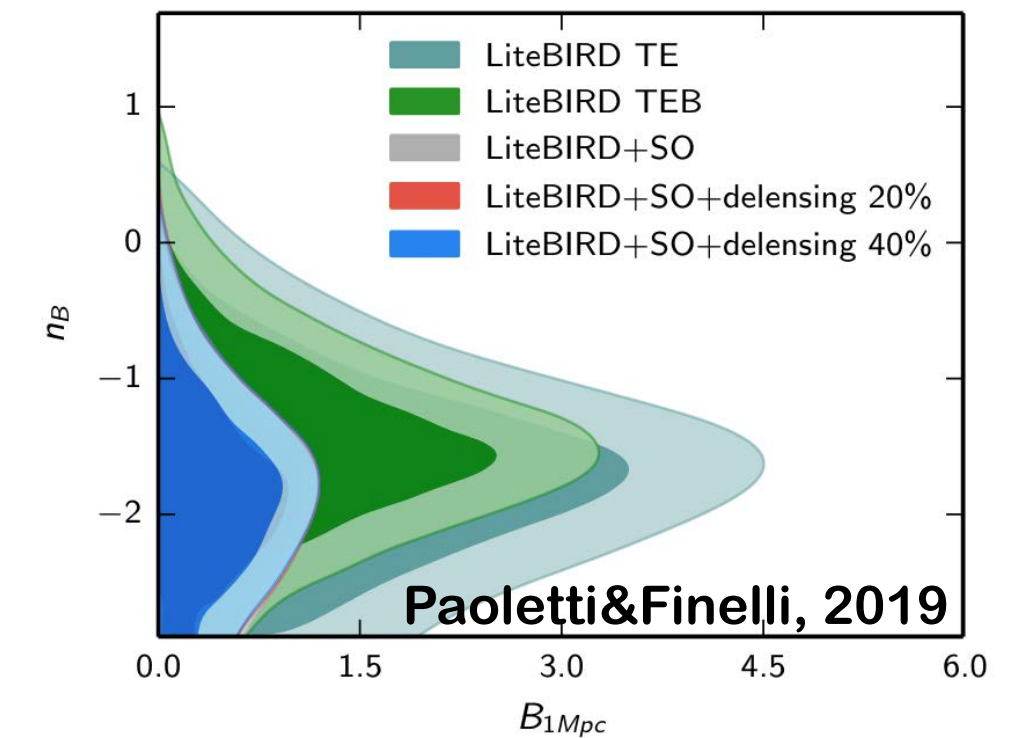
- Stronger significance
- Cleaner model characterisation



ψ experiments	σ_r
S4-Wide + SO-SAT + ASO-SAT	0.0110
S4-Wide + LiteBIRD	0.0077
S4-Deep + S4D-SAT	0.0064
Namikawa+, 2025 Combined	0.0047

Lensing potential as a matter tracer from small-scale CMB improve r constraints:

- With efficient delensing
- By allowing alternative analysis methods (e.g., Namikawa&Sherwin, 2023)

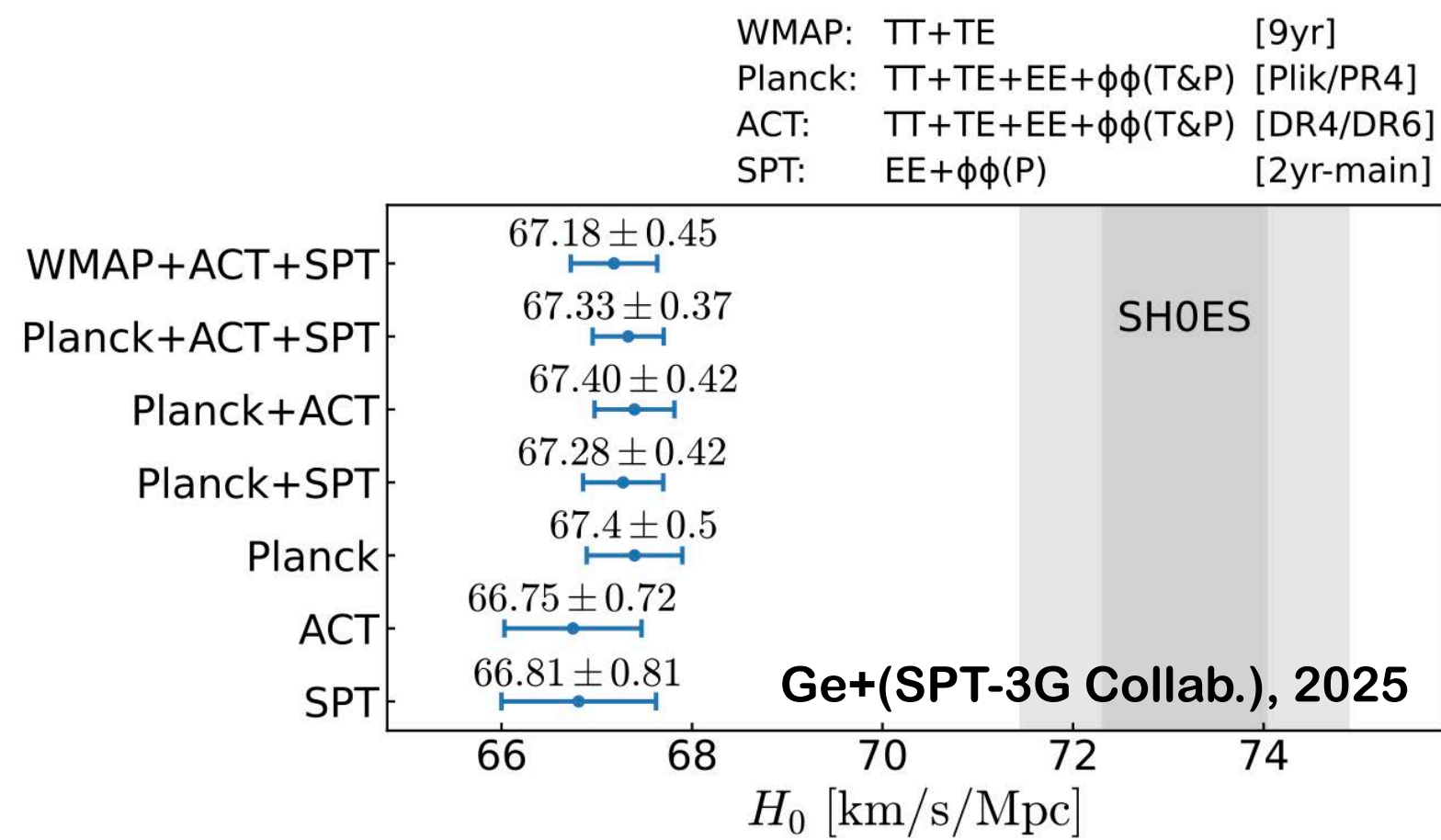


Dataset ($A_L = 1$)	Fields	τ	$\beta_E^2 [10^{-2}]$	$\beta_V^2 [10^{-3}]$	$r [10^{-3}]$
LB w/q small scale	TEB	$0.0543^{+0.0019}_{-0.0021}$	< 9.7	< 1.1	< 0.4
	TEBV	0.0544 ± 0.0020	< 0.5	< 1.1	< 0.4
LB+ CMB-S4	TEB	$0.0544^{+0.0017}_{-0.0020}$	< 5.1	< 0.9	< 0.4
	TEBV	0.0544 ± 0.0019	< 0.5	< 0.8	< 0.4

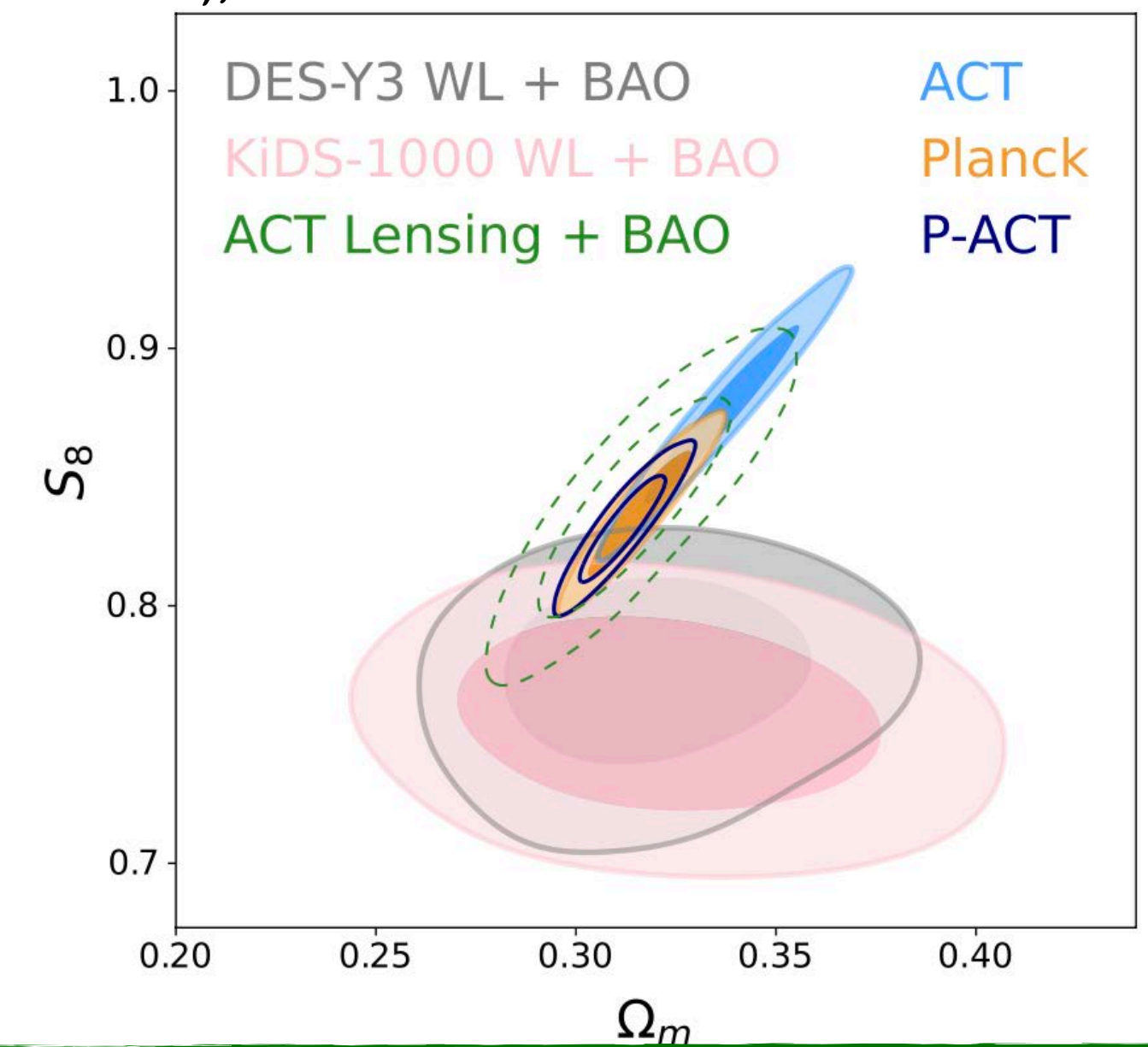
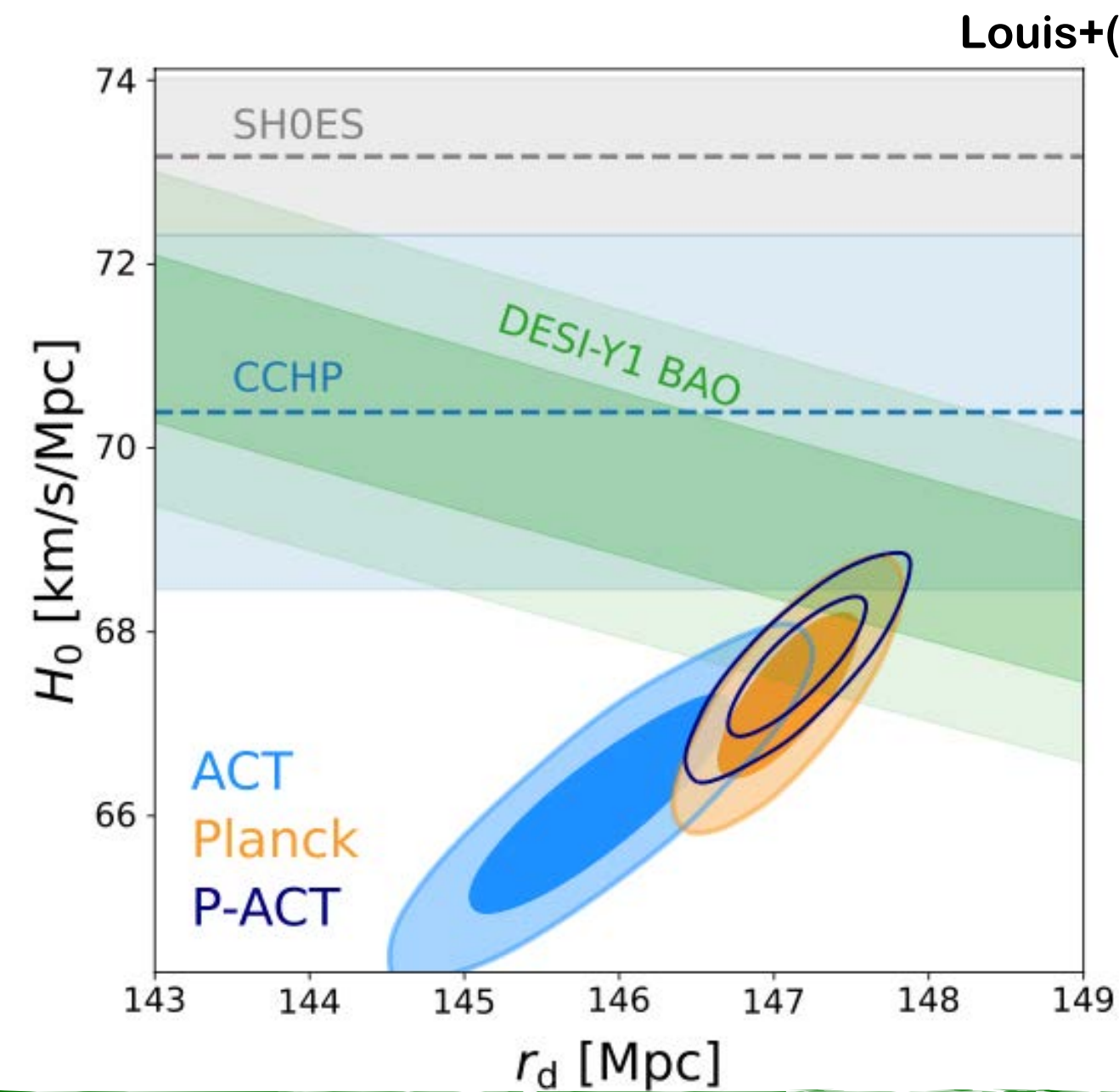
Raffuzzi+, 2024

Combination of different imprints on different scales and fields improves constraints on physical model parameters (e.g., Paoletti&Finelli, 2019 on Primordial Magnetic Fields; Raffuzzi+, 2024 on Generalised Faraday Effect)

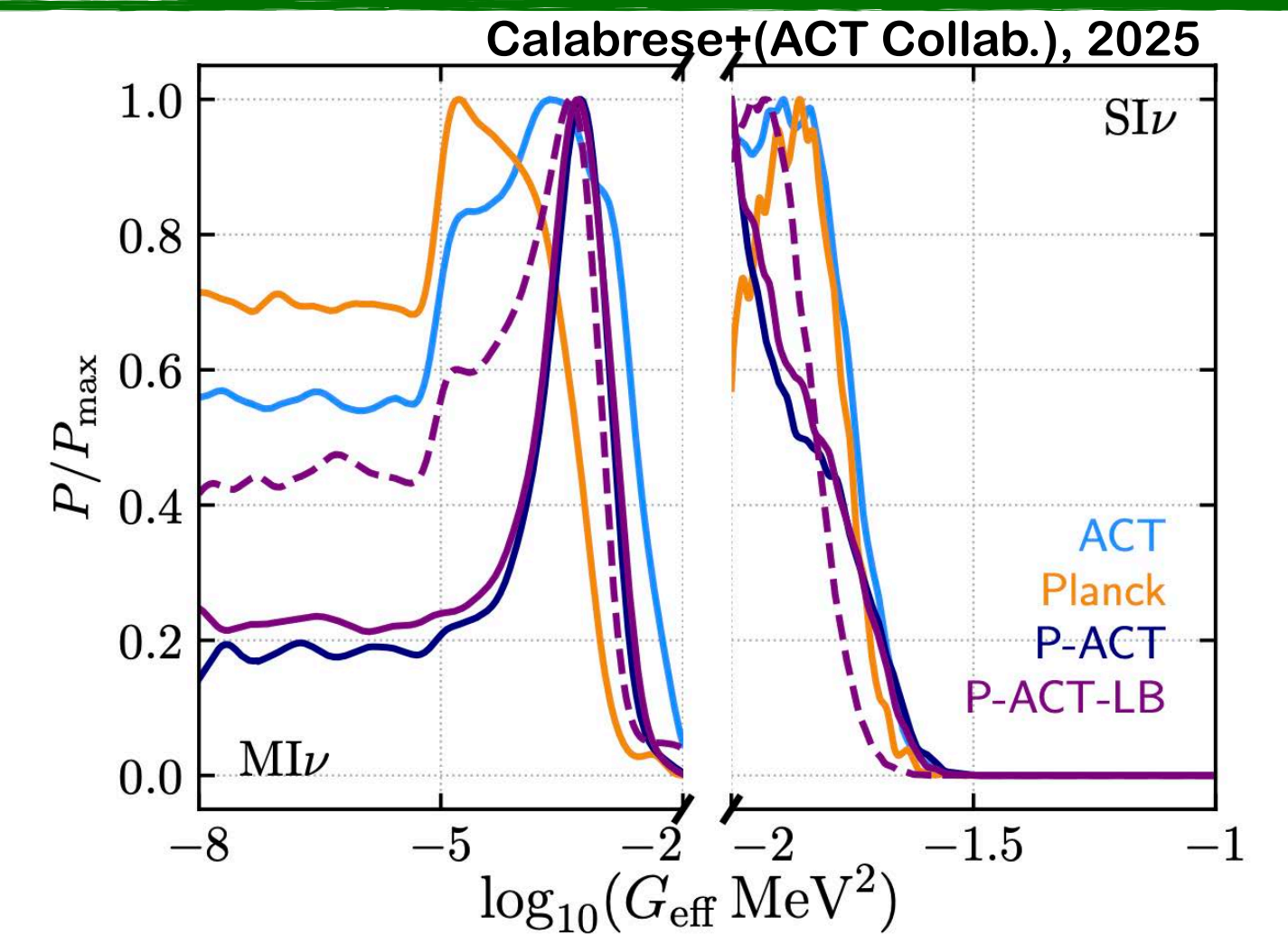
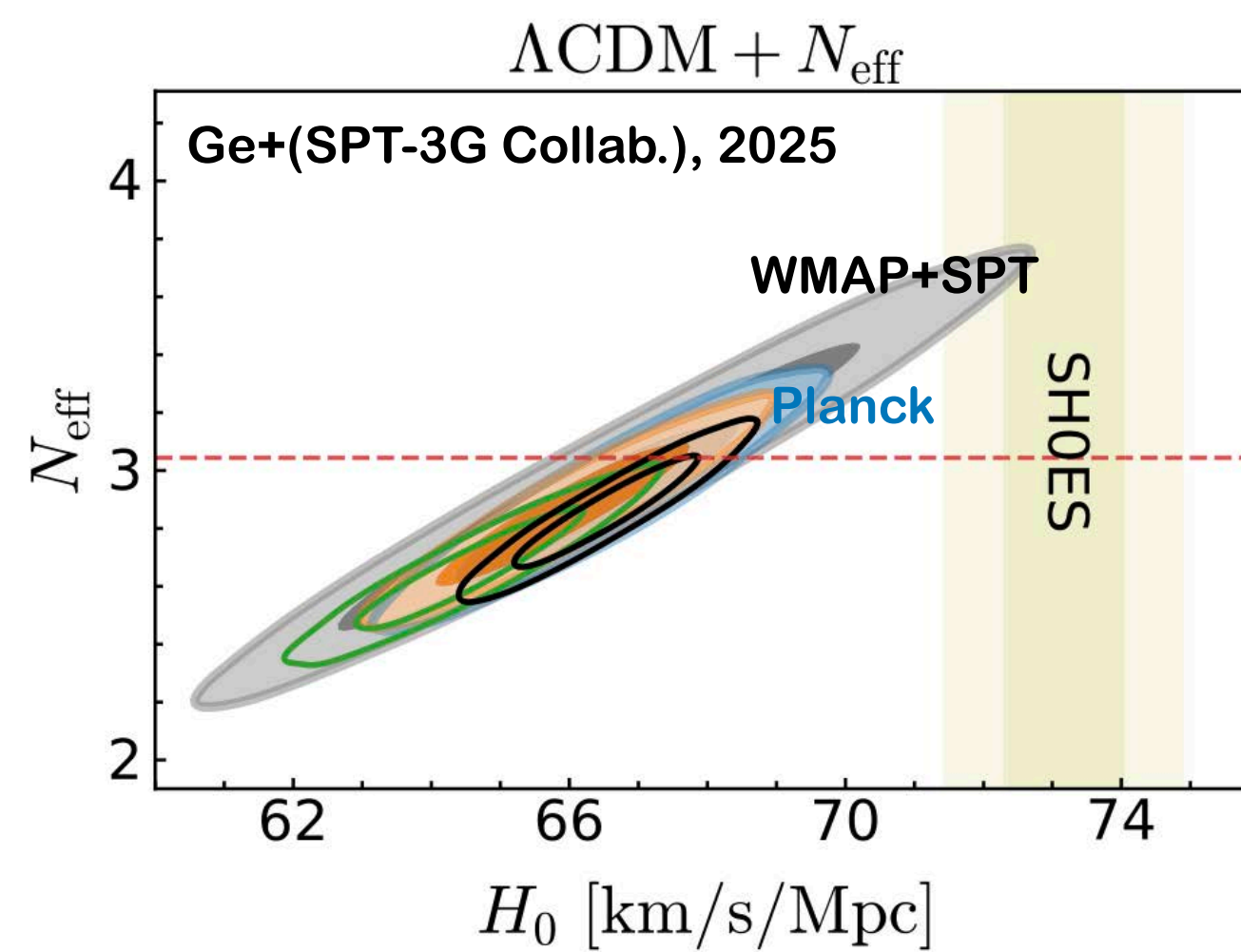
Cosmic concordance



Independent and comparable estimates are key in understanding tensions

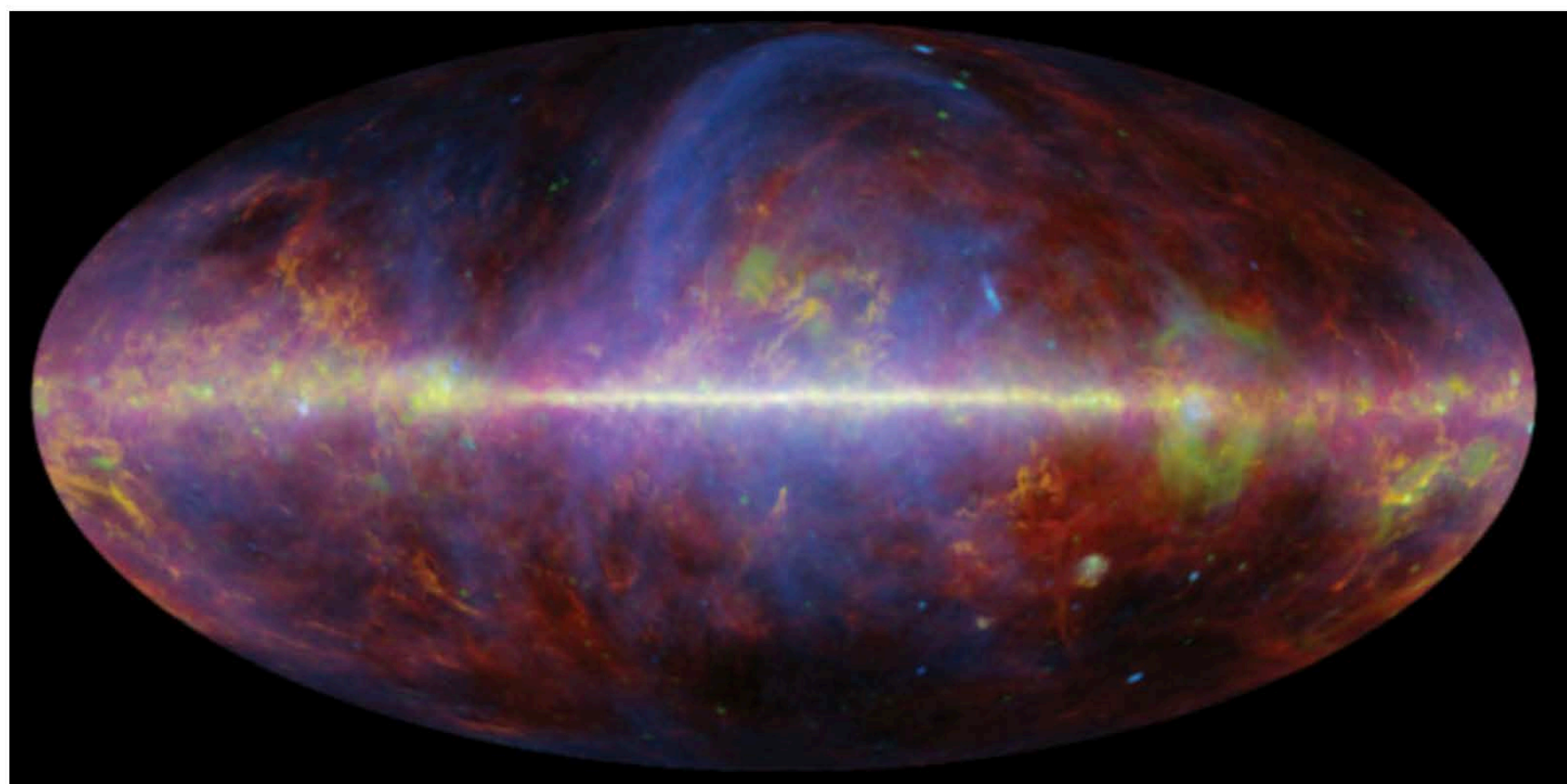


Independent and comparable estimates are key in strengthening confidence in model constraints



Cosmo+astrophysics

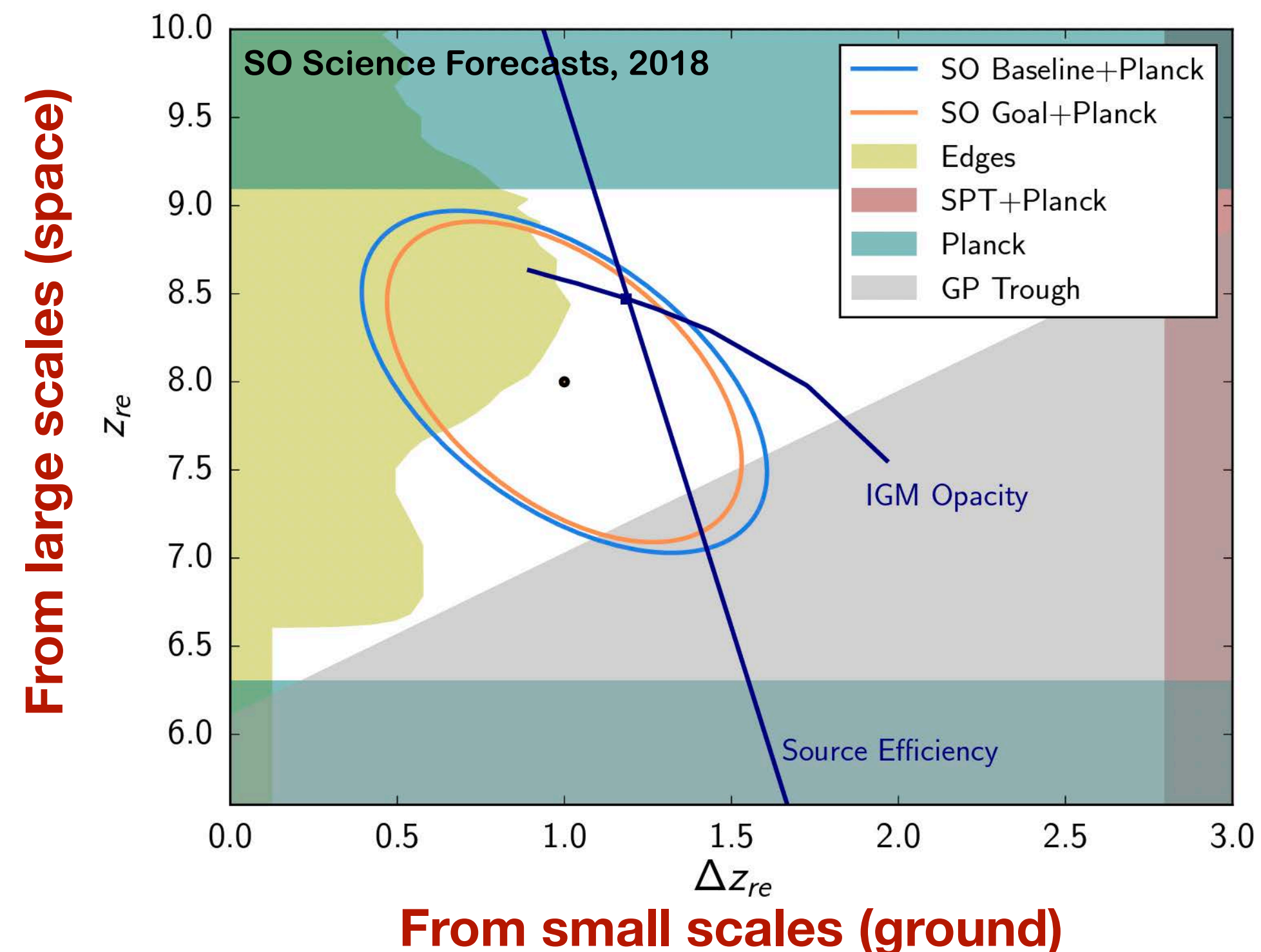
Panexperiment Galactic Science Group



Composite image of Galactic emission as measured by the Planck satellite. Different colors represent dust, synchrotron, free-free, and carbon monoxide line emission. Credit: ESA/NASA/JPL-Caltech.

Our mission

The purpose of the Pan-Experiment Galactic Science Group is to pool expertise from researchers interested in using cosmic microwave background experiments to study the Milky Way. The group operates as a virtual meeting space for scientists that work across various



“Deep and increasing connection between cosmology and the rest of astrophysics. Cosmological probes are invariably intertwined with their astrophysical context. These connections are often couched as “systematic uncertainties,” which ignores the synergistic opportunities that come with the co-development of different areas of the field”

<https://doi.org/10.17226/26141>.

Some discussion points

- Are there any science targets not yet identified which could benefit from space/ground complementarity?
- Overlapping communities: how to make sure analyses and results are “uncorrelated”?
- Shall we identify new (governance/computational/infrastructural/funding) tools and platforms to boost collaborative efforts?
- Public products: enough to guarantee reproducibility, cross-checks, info sharing?
- Data format/simulations/analysis tools: do we need coordinated efforts?

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